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On W_2 -Curvature Tensor $N(k)$ -Quasi Einstein Manifolds

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ABSTRACT

We consider $N(k)$ -quasi Einstein manifolds satisfying the conditions $R(\xi, X).W_2 = 0$, $W_2(\xi, X).S = 0$, $P(\xi, X).W_2 = 0$, where W_2 and P denote the W_2 -curvature tensor and projective curvature tensor, respectively.

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1. Introduction

The notion of a quasi Einstein manifold was introduced by M. C. Chaki in [1]. A non flat n -dimensional Riemannian manifold (M, g) is said to be a quasi Einstein manifold if its Ricci tensor S satisfies

$$(1.1) \quad S(X, Y) = ag(X, Y) + b\eta(X)\eta(Y), \quad \forall X, Y \in TM$$

for some smooth functions a and $b \neq 0$, where η is a non zero 1-forms such that

$$(1.2) \quad g(X, \xi) = \eta(X), \quad g(\xi, \xi) = \eta(\xi) = 1,$$

for the associated vector field ξ . The 1-form η is called the associated 1-form and the unit vector field ξ is called the generator of the manifold. If $b \neq 0$ then the manifold reduced to an Einstein manifold. For more details about quasi Einstein manifolds see also ([2]-[6]).

If the generator ξ belongs to k -nullity distribution $N(k)$ then the quasi Einstein manifold is called as an $N(k)$ -quasi Einstein manifold [12]. In [12], it was shown that a conformally flat quasi Einstein manifold is an $N(k)$ -quasi Einstein manifold and in particular a 3-dimensional quasi Einstein manifold is an $N(k)$ -quasi Einstein manifold. The derivation conditions $R(\xi, X).R = 0$ and $R(\xi, X).S = 0$ were also studied in [12], where R and S denote the curvature and Ricci

tensor, respectively. In [9], it was proved that in an n-dimensional $N(k)$ -quasi Einstein manifold $k = \frac{a+b}{n-1}$. In [7], derivation conditions $R(\xi, X). \rho = 0, \rho(\xi, X). S = 0$ and $\rho(\xi, X). \rho = 0$ were studied where ρ is the projective curvature tensor, also it physical examples of $N(k)$ -quasi Einstein manifolds were given. The derivation conditions $R(\xi, X). C = 0, R(\xi, X). \tilde{C} = 0$, studied in [8], where C and \tilde{C} denote the conformal curvature tensor and quasi conformal curvature tensor, respectively. In this paper, we consider $N(k)$ -quasi Einstein manifolds satisfying the conditions $R(\xi, X). W_2 = 0, W_2(\xi, X). S = 0, P(\xi, X). W_2 = 0$, where W_2 and P denote the W_2 -curvature tensor and the projective curvature tensor, respectively.

2. $N(k)$ -quasi Einstein manifolds

From (1.1) and (1.2) we obtain

$$(2.1) \quad S(X, \xi) = (a + b)\eta(X)$$

$$(2.2) \quad r = na + b,$$

where r is the scalar curvature of M .

The Ricci operator Q of a Riemannian manifold (M, g) is defined by

$$S(X, Y) = g(QX, Y).$$

If (M, g) is a quasi Einstein manifold [1], its Ricci operator satisfies

$$(2.3) \quad Q = aI + b\eta \otimes \xi.$$

Let R denote the Riemannian curvature tensor of a Riemannian manifold M . The k -nullity distribution $N(k)$ [11], of a Riemannian manifold defined by

$$N(k): p \rightarrow N_p(k) = \{Z \in T_p M \mid R(X, Y)Z = k\{g(Y, Z)X - g(X, Z)Y\}\}$$

for all $X, Y \in TM^n$, where k is some smooth function. In a quasi Einstein manifold M , if the generator ξ belongs to some k -nullity distribution $N(k)$, then is said to be an $N(k)$ -quasi Einstein manifold [12].

Lemma2.1. [9] In an n-dimensional $N(k)$ -quasi Einstein manifold it follows that

$$(2.4) \quad k = \frac{a+b}{n-1}.$$

Let (M^n, g) be an $N(k)$ -quasi Einstein manifold. Then, we have [12]

$$(2.5) \quad R(Y, Z)\xi = \frac{a+b}{n-1} \{\eta(Z)Y - \eta(Y)Z\}.$$

The equation (2.5) is equivalent to

$$(2.6) \quad R(\xi, Y)Z = \frac{a+b}{n-1} \{g(Y, Z)\xi - \eta(Z)Y\} = -R(Y, \xi)Z.$$

In [7], we view the following physical examples of $N(k)$ -quasi Einstein manifolds.

Example2.2. [7] A conformally flat perfect fluid space-time (M^4, g) satisfying Einstein's equation without cosmological constant is an $N\left(\frac{k(3\sigma + p)}{6}\right)$ -quasi Einstein manifold.

Example2.3. [7] A conformally flat perfect fluid spacetime (M^4, g) satisfying Einstein's equation with cosmological constant is an $N\left(\frac{\lambda}{3} + \frac{k(3\sigma + p)}{6}\right)$ -quasi Einstein manifold.

where σ is the energy density and p is the isotropic pressure of the fluid.

3. The W_2 -curvature tensor of an $N(k)$ -quasi Einstein manifold

In [10], Pokhariyal and Mishra have defined a new curvature tensor

$$(3.1) \quad W_2(X, Y, Z, U) = R(X, Y, Z, U) + \frac{1}{n-1} [g(X, Z)S(Y, U) - g(Y, Z)S(X, U)],$$

where S is a Ricci tensor of type (0,2).

Proposition 3.1. In an n -dimensional $N(k)$ -quasi Einstein manifold M , the W_2 -curvature tensor satisfies

$$(3.2) \quad \eta(W_2(X, Y)Z) = 0,$$

for all vector fields X, Y, Z on M .

Proof. From (1.1), (2.5), (2.6) and (3.1) the Eq. (3.2) follow easily. □

Theorem 3.2. Let M be an $N(k)$ -quasi Einstein manifold. Then M satisfies the condition $R(\xi, X).W_2 = 0$ if and only if $a + b = 0$ or $W_2 = 0$.

Proof. Assume that M is an n -dimensional $N(k)$ -quasi Einstein manifold satisfying the condition $R(\xi, X).W_2 = 0$. Since M satisfies the condition $R(\xi, X).W_2 = 0$ we can write

$$0 = R(\xi, X)W_2(Y, Z)U - W_2(R(\xi, X)Y, Z)U \\ - W_2(Y, R(\xi, X)Z)U - W_2(Y, Z)R(\xi, X)U$$

for all vector fields X, Y, Z, U on M .

Using (2.6), in above equation we get

$$0 = \frac{a+b}{n-1} \{W_2(Y, Z, U, X)\xi - \eta(W_2(Y, Z)U)X \\ - g(X, Y)W_2(\xi, Z)U + \eta(Y)W_2(X, Z)U \\ - g(X, Z)W_2(Y, \xi)U + \eta(Z)W_2(Y, X)U \\ - g(X, U)W_2(Y, Z)\xi + \eta(U)W_2(Y, Z)X \}$$

which implies either $a + b = 0$ or

$$0 = W_2(Y, Z, U, X)\xi - \eta(W_2(Y, Z)U)X \\ - g(X, Y)W_2(\xi, Z)U + \eta(Y)W_2(X, Z)U \\ - g(X, Z)W_2(Y, \xi)U + \eta(Z)W_2(Y, X)U \\ - g(X, U)W_2(Y, Z)\xi + \eta(U)W_2(Y, Z)X.$$

Taking the inner product of above equation with ξ we obtain

$$(3.3) \quad 0 = W_2(Y, Z, U, X) - \eta(W_2(Y, Z)U)\eta(X) \\ - g(X, Y)\eta(W_2(\xi, Z)U) + \eta(Y)\eta(W_2(X, Z)U) \\ - g(X, Z)\eta(W_2(Y, \xi)U) + \eta(Z)\eta(W_2(Y, X)U) \\ - g(X, U)\eta(W_2(Y, Z)\xi) + \eta(U)\eta(W_2(Y, Z)X).$$

Using (3.2) into (3.3) we obtain

$$W_2(Y, Z, W, X) = 0.$$

The converse statement is trivial. This completes the proof of the theorem. □

Next, we have the following theorem

Theorem 3.3. Let M be an $N(k)$ -quasi Einstein manifold. Then M satisfies the condition $W_2(\xi, X).S = 0$ if and only if $a = 0$.

Proof. Since $W_2(\xi, X).S = 0$, we have

$$(3.4) \quad 0 = S(W_2(\xi, X)Y, Z) + S(Y, W_2(\xi, X)Z).$$

In view of (1.1), (3.1) and (3.2) in (3.4) we have

$$(3.5) \quad 0 = \frac{ab}{n-1} \{g(X, Y)\eta(Z) + g(X, Z)\eta(Y) - 2\eta(X)\eta(Y)\mu(Z)\}.$$

From (3.5) by a contraction, we obtain

$$(3.6) \quad ab = 0.$$

Since $b \neq 0$, then from (3.6) we have $a = 0$. The converse statement is trivial. This completes the proof of the theorem. \square

Let (M^n, g) be a Riemannian manifold. The projective curvature tensor [13], is defined by

$$(3.7) \quad P(X, Y)Z = R(X, Y)Z - \frac{1}{n-1} [S(Y, Z)X - S(X, Z)Y],$$

If P be a projective curvature tensor in an n -dimensional $N(k)$ -quasi Einstein manifold, we have [7]

$$(3.8) \quad P(\xi, X)Y = \frac{b}{n-1} [g(X, Y)\xi - \eta(X)\eta(Y)\xi].$$

Theorem 3.4. Let M be an $N(k)$ -quasi Einstein manifold. Then M satisfies the condition $P(\xi, X).W_2 = 0$ if and only if $W_2 = 0$.

Proof. Assume that M is an n -dimensional $N(k)$ -quasi Einstein manifold satisfying the condition $P(\xi, X).W_2 = 0$. Since M satisfies the condition $P(\xi, X).W_2 = 0$ we can write

$$(3.9) \quad 0 = P(\xi, X)W_2(Y, Z)U - W_2(P(\xi, X)Y, Z)U \\ - W_2(Y, P(\xi, X)Z)U - W_2(Y, Z)P(\xi, X)U$$

for all vector fields X, Y, Z, U on M .

Using (3.8), in (3.9) equation we get

$$0 = \frac{b}{n-1} \{W_2(Y, Z, U, X)\xi - \eta(W_2(Y, Z)U)\eta(X)\xi \\ - g(X, Y)W_2(\xi, Z)U + \eta(X)\eta(Y)W_2(\xi, Z)U \\ - g(X, Z)W_2(Y, \xi)U + \eta(Z)\eta(X)W_2(Y, \xi)U \\ - g(X, U)W_2(Y, Z)\xi + \eta(X)\eta(U)W_2(Y, Z)\xi\}.$$

Since $b \neq 0$, then

$$\begin{aligned}
 (3.10) \quad 0 &= W_2(Y, Z, U, X)\xi - \eta(W_2(Y, Z)U)\eta(X)\xi \\
 &\quad - g(X, Y)W_2(\xi, Z)U + \eta(X)\eta(Y)W_2(\xi, Z)U \\
 &\quad - g(X, Z)W_2(Y, \xi)U + \eta(Z)\eta(X)W_2(Y, \xi)U \\
 &\quad - g(X, U)W_2(Y, Z)\xi + \eta(X)\eta(U)W_2(Y, Z)\xi.
 \end{aligned}$$

Taking the inner product of (3.10) with ξ we obtain

$$\begin{aligned}
 (3.11) \quad 0 &= W_2(Y, Z, U, X) - \eta(W_2(Y, Z)U)\eta(X) \\
 &\quad - g(X, Y)\eta(W_2(\xi, Z)U) + \eta(X)\eta(Y)\eta(W_2(\xi, Z)U) \\
 &\quad - g(X, Z)\eta(W_2(Y, \xi)U) + \eta(Z)\eta(X)\eta(W_2(Y, \xi)U) \\
 &\quad - g(X, U)\eta(W_2(Y, Z)\xi) + \eta(X)\eta(U)\eta(W_2(Y, Z)\xi).
 \end{aligned}$$

Using (3.2) into (3.3) we obtain

$$W_2(Y, Z, W, X) = 0.$$

The converse statement is trivial. This completes the proof of the theorem. \square

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